

# Magnetar giant flare in NGC 253 seen by Fermi-GBM

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## Rapid spectral variability of a giant flare from a magnetar in NGC 253

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2049 Accesses | 1 Citations | 201 Altmetric | Metrics

### Abstract

Magnetars are neutron stars with extremely strong magnetic fields ( $10^{13}$  to  $10^{15}$  gauss)<sup>1,2</sup>, which episodically emit X-ray bursts approximately 100 milliseconds long and with energies of  $10^{40}$  to  $10^{41}$  erg. Occasionally, they also produce extremely bright and energetic giant flares, which begin with a short (roughly 0.2 seconds), intense flash, followed by fainter, longer-lasting emission that is modulated by the spin period of the magnetar<sup>3,4</sup> (typically 2 to 12 seconds). Over the past 40 years, only three such flares have been observed in our local group of galaxies<sup>3,4,5,6</sup>, and in all cases the extreme intensity of the flares caused the detectors to saturate. It has been proposed that extragalactic giant flares are probably a subset<sup>7,8,9,10,11</sup> of short  $\gamma$ -ray bursts, given that the sensitivity of current instrumentation prevents us from detecting the pulsating tail, whereas the initial bright flash is readily observable out to distances of around 10 to 20 million parsecs. Here we report X-ray and  $\gamma$ -ray observations of the  $\gamma$ -ray burst GRB 200415A, which has a rapid onset, very fast time variability, flat spectra and substantial sub-millisecond spectral evolution. These attributes match well with those expected for a giant flare from an extragalactic magnetar<sup>12</sup>, given that GRB 200415A is directionally associated<sup>13</sup> with the galaxy NGC 253 (roughly 3.5 million parsecs away). The detection of three mega-electronvolt photons provides evidence for the relativistic motion of the emitting plasma. Radiation from such rapidly moving gas around a rotating magnetar may have generated the rapid spectral evolution that we observe.



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12–23 July 2021



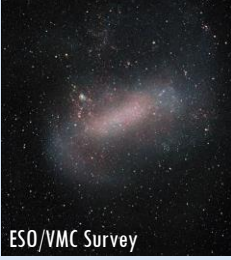




# Magnetar Giant Flares

## ■ Magnetars

- Strongly magnetized ( $\sim 10^{15}$  G) **neutron stars**, period  $\sim 0.1-10$  s, luminosity  $L_x \sim 10^{31-36}$  erg/s
  - Never observed at  $E > 100$  MeV (before April 15<sup>th</sup>, 2020!)
- **Recurrent, short duration bursts** ( $10^{36-41}$  erg/s), with period  $\sim 0.1$  s
- **Long lasting bursts** ( $> 10^{36}$  erg/s) lasting years (total energy  $E \sim 10^{41-43}$  erg)
- **Giant Flares**  $L \sim 10^{44-47}$  erg/s, lasting  $\sim 0.1$  s, followed by soft X-ray emission lasting hundreds of seconds



Burns+ApJL (2021)

March 5, 1979	November 3, 2005	February 1, 2007	February 22, 2007	April 15, 2020
Large Magellanic Cloud	M82 (Cigar Galaxy)	M31 (Andromeda)	M83 (Pinwheel)	NGC 253 (Sculptor)
				
163 kly	11.4 Mly	2.5 Mly	14.7 Mly	11.4 Mly
$0.7 \times 10^{44}$ erg	$5.3 \times 10^{46}$ erg	$1.6 \times 10^{45}$ erg	$6.2 \times 10^{45}$ erg	$1.3 \times 10^{46}$ erg*

- **Bright transient** triggered the Inter-Planetary Network (IPN) and ASIM on **April 15<sup>th</sup>, 2020**

- **Fermi Gamma-ray Burst Monitor (GBM) Trigger**

at 08:48:05.56 UTC

*The Fermi GBM team, GCN 27579 (2020)*

*E. Bissaldi, et al., GCN 27587 (2020)*

*O. Roberts et al., Nature Vol.589, 207 (2021)*

- Triangulated by IPN to a 17 sq. arcmin region centered at

RA. 11.88°, Dec. -25.26° (J2000), overlapping with

**NGC 253 (Sculptor Galaxy)**, active star-burst spiral galaxy

at a distance of 3.5 Mpc (11.4 Mly)

*D. Svinkin, et al., GCN 27585 (2021)*

*D. Svinkin, et al., Nature Vol.589, 211 (2021)*

- Chance coincidence with NGC 253: **1 in 230,000**

[Talk by M. Negro, ID.792, PoS(ICRC2021)630]

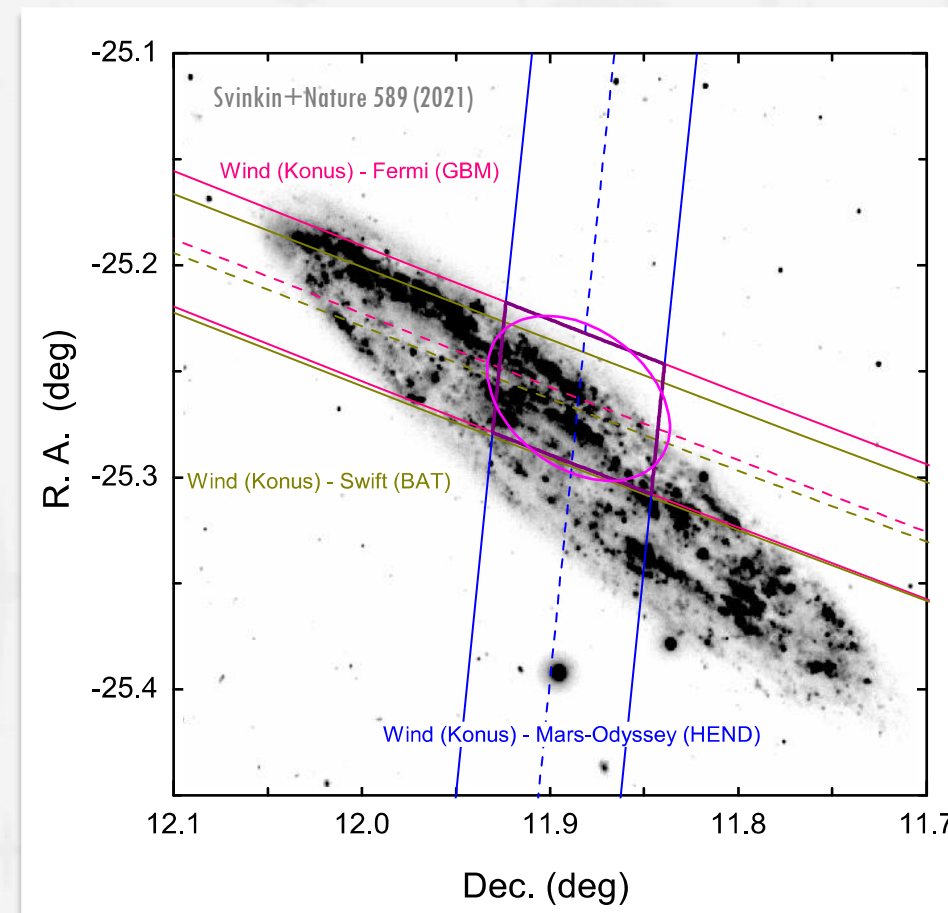
*E. Burns et al., ApJL 907 L28 (2021)*

- **High Energy (GeV) emission** [[Highlight Talk by N.Di Lalla, ID.1480, PoS\(ICRC2021\)019](#)]

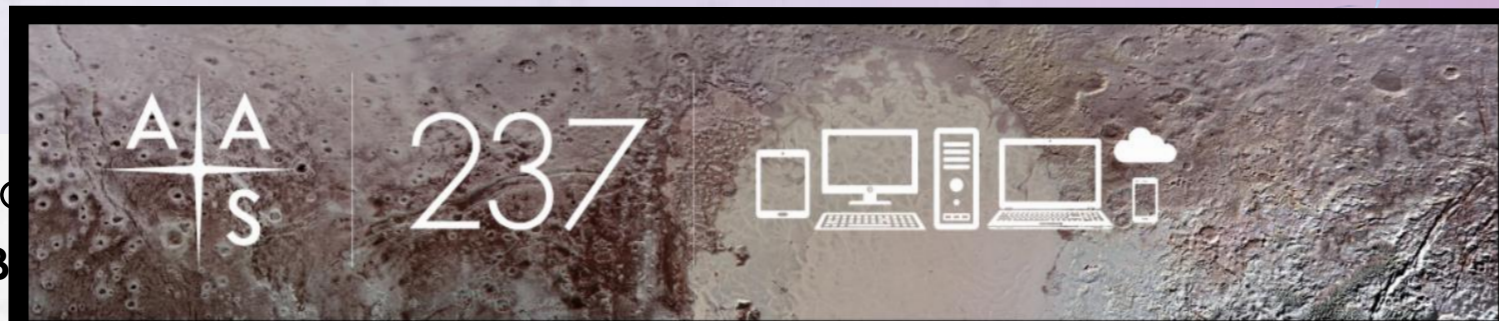
- **Fermi Large Area Telescope (LAT) detection**

*N. Omodei, et al. GCN 27586, GCN 27597 (2020)*

*M. Ajello, et al., Nature Astronomy, Vol.5, 385 (2021)*



- **Bright transient** triggered the Inter-Planetary Network (IPN)
  - **Fermi Gamma-ray Burst Monitor (GBM)** detected GRB 200415A at 08:48:05.56 UTC  
*The Fermi GBM team, GCN 27579 (2020)*  
*E. Bissaldi, et al., GCN 27587 (2020)*  
*O. Roberts et al., Nature Vol.589, 207 (2021)*
  - Triangulated by IPN to a 17 sq. arcmin area, RA. 11.88°, Dec. -25.26° (J2000), over **NGC 253 (Sculptor Galaxy)**, active star-forming galaxy at a distance of 3.5 Mpc (11.4 Mly)  
*D. Svinkin, et al., GCN 27585 (2021)*  
*D. Svinkin, et al., Nature Vol.589, 211 (2021)*
  
- **Chance coincidence with NGC 253**  
 [Talk by M. Negro, ID.792, PoS(ICRC2021)630]  
*E. Burns et al., ApJL 907 L28 (2021)*
  
- **High Energy (GeV) emission** [Highlight Talk by M. Ajello]
  - **Fermi Large Area Telescope (LAT)** detected a high-energy emission  
*N. Omodei, et al. GCN 27586, GCN 27597 (2020)*  
*M. Ajello, et al., Nature Astronomy, Vol.5, 385 (2021)*



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**SESSION (Press Conference) Bursting Magnetars**

Itinerary

January 13, 2021

Virtual Meeting - Press

DESCRIPTION

Bursting Magnetars

A Masquerading Magnetar in the Sculptor Galaxy  
 Kevin Hurley (University of California, Berkeley)

GRB 200415A: Gamma-Ray Burst or Magnetar Giant Flare?  
 Oliver Roberts (Universities Space Research Association)

High-Energy Emission from a Magnetar Giant Flare  
 Nicola Omodei (Stanford University)

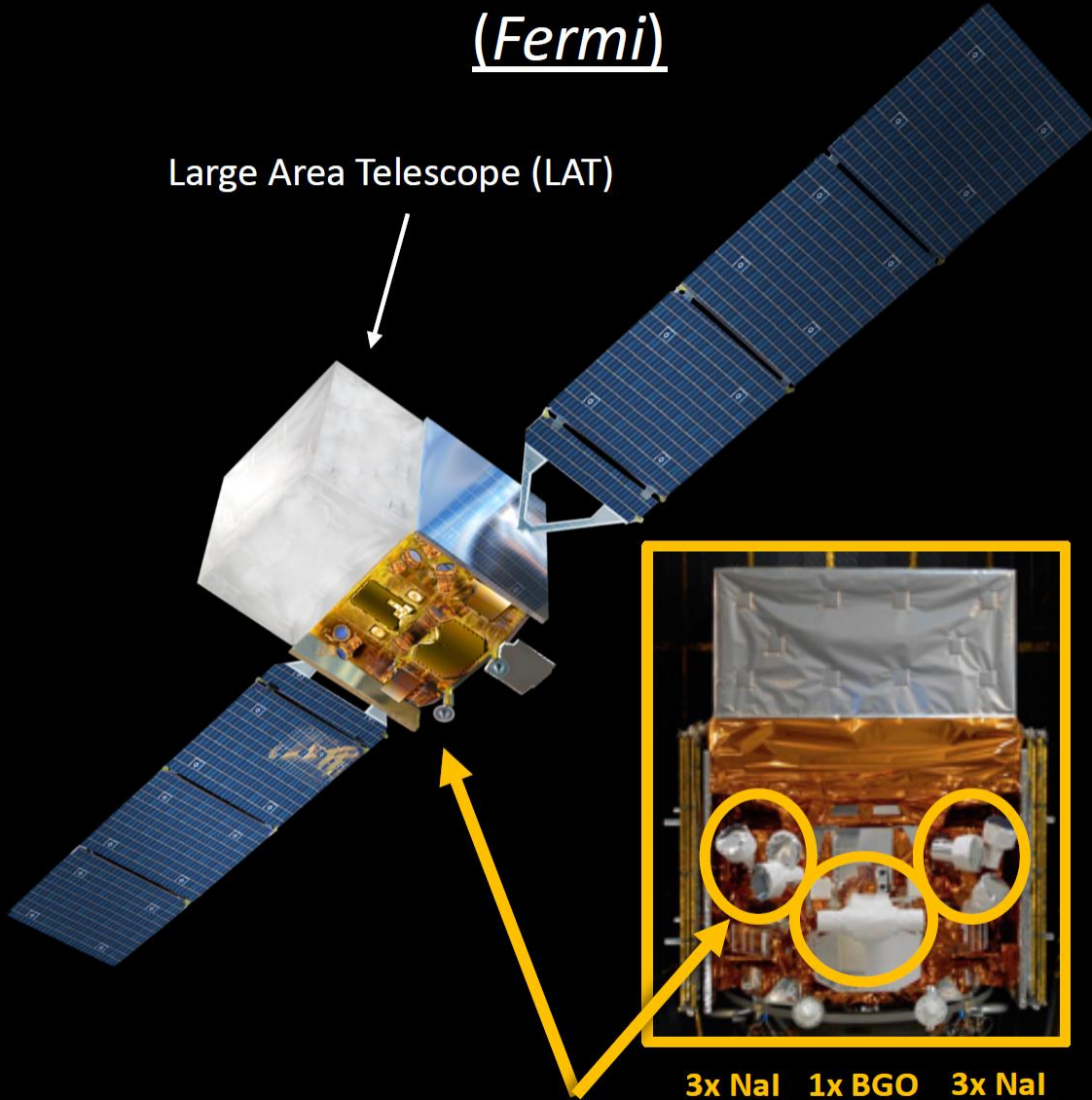
Extragalactic Magnetar Giant Flares Are a Source of Gamma-Ray Bursts  
 Eric Burns (Louisiana State University)

**AAS Press Briefing:**

**<https://www.youtube.com/watch?v=LjroNW7D-E4>**

## Fermi Gamma-ray Space Telescope (Fermi)

Large Area Telescope (LAT)



3x NaI 1x BGO 3x NaI

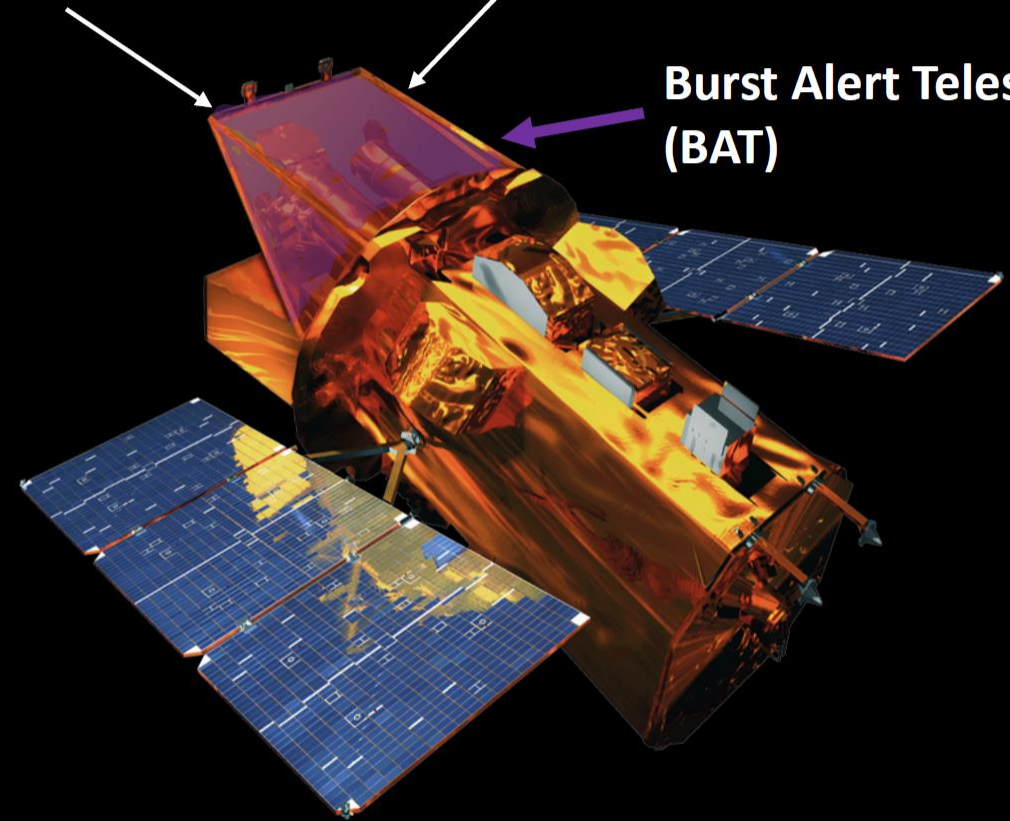
**Gamma-ray Burst Monitor (GBM)**

## Neil Gehrels Swift Observatory (Swift)

X-Ray Telescope (XRT)

Ultraviolet/Optical Telescope (UVOT)

**Burst Alert Telescope (BAT)**

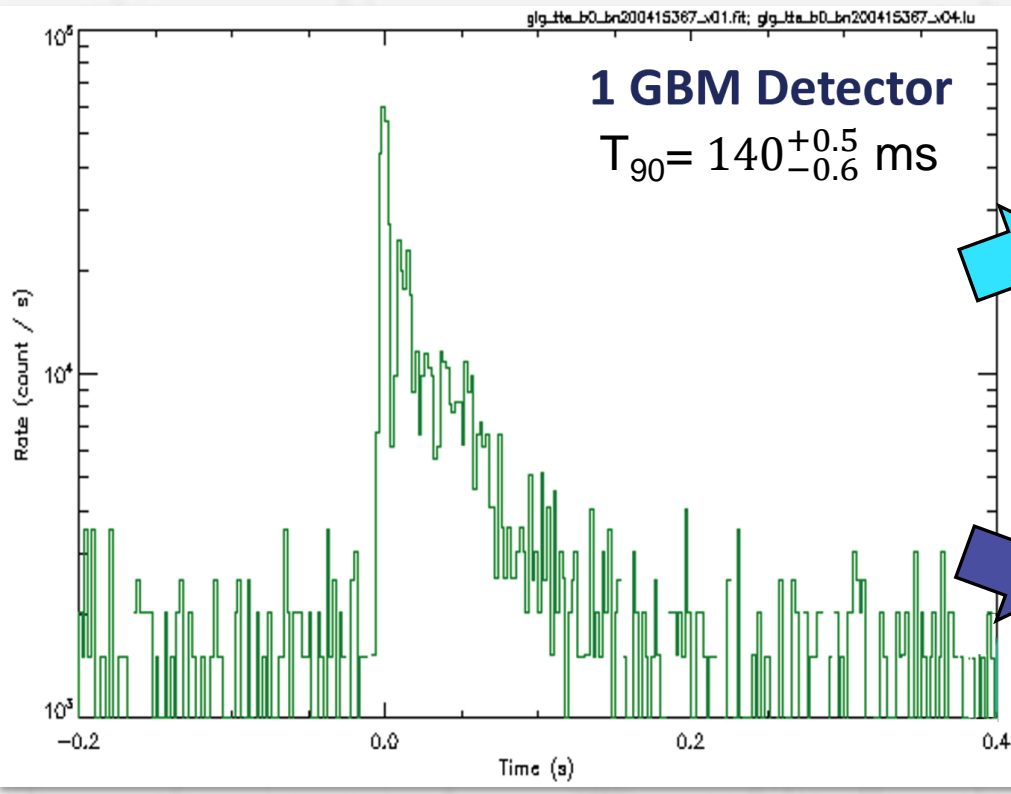


GRB 200415A detected offline using the Gamma-ray Urgent Archiver for Novel Opportunities (**GUANO**), a BAT pipeline to search for transients coincident with Gravitational waves.

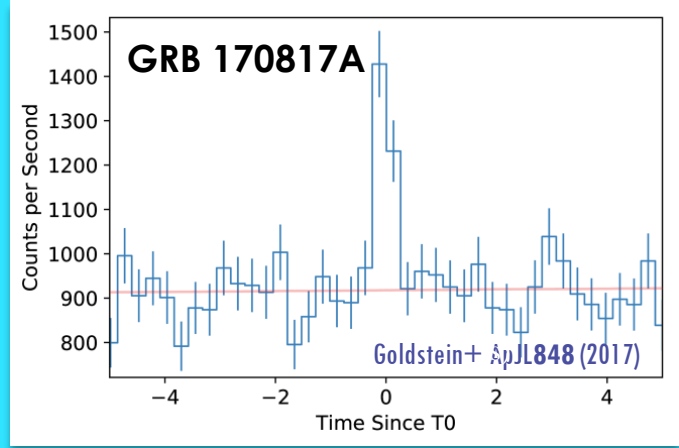
# GRB 200415A



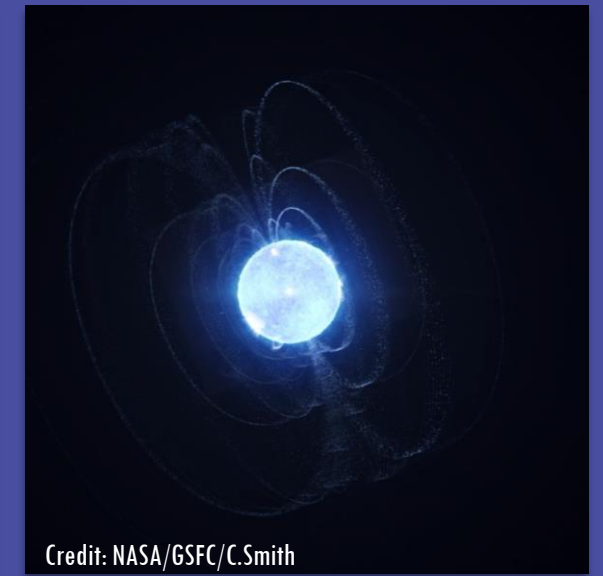
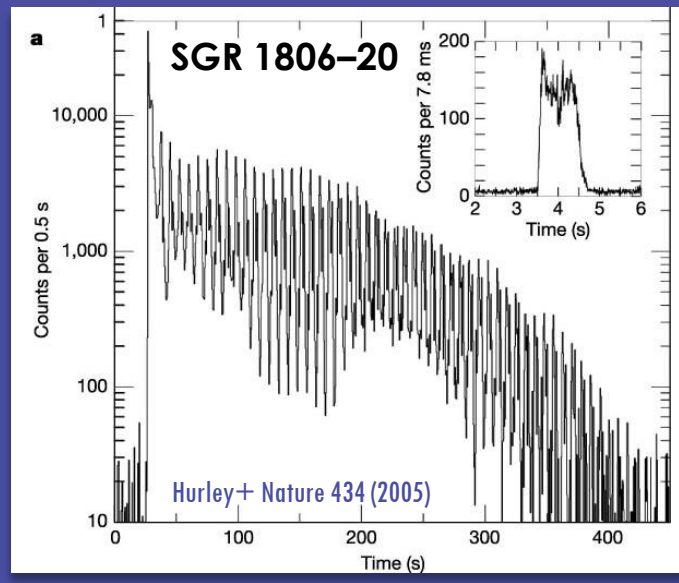
## Short Gamma-ray Burst (sGRB)



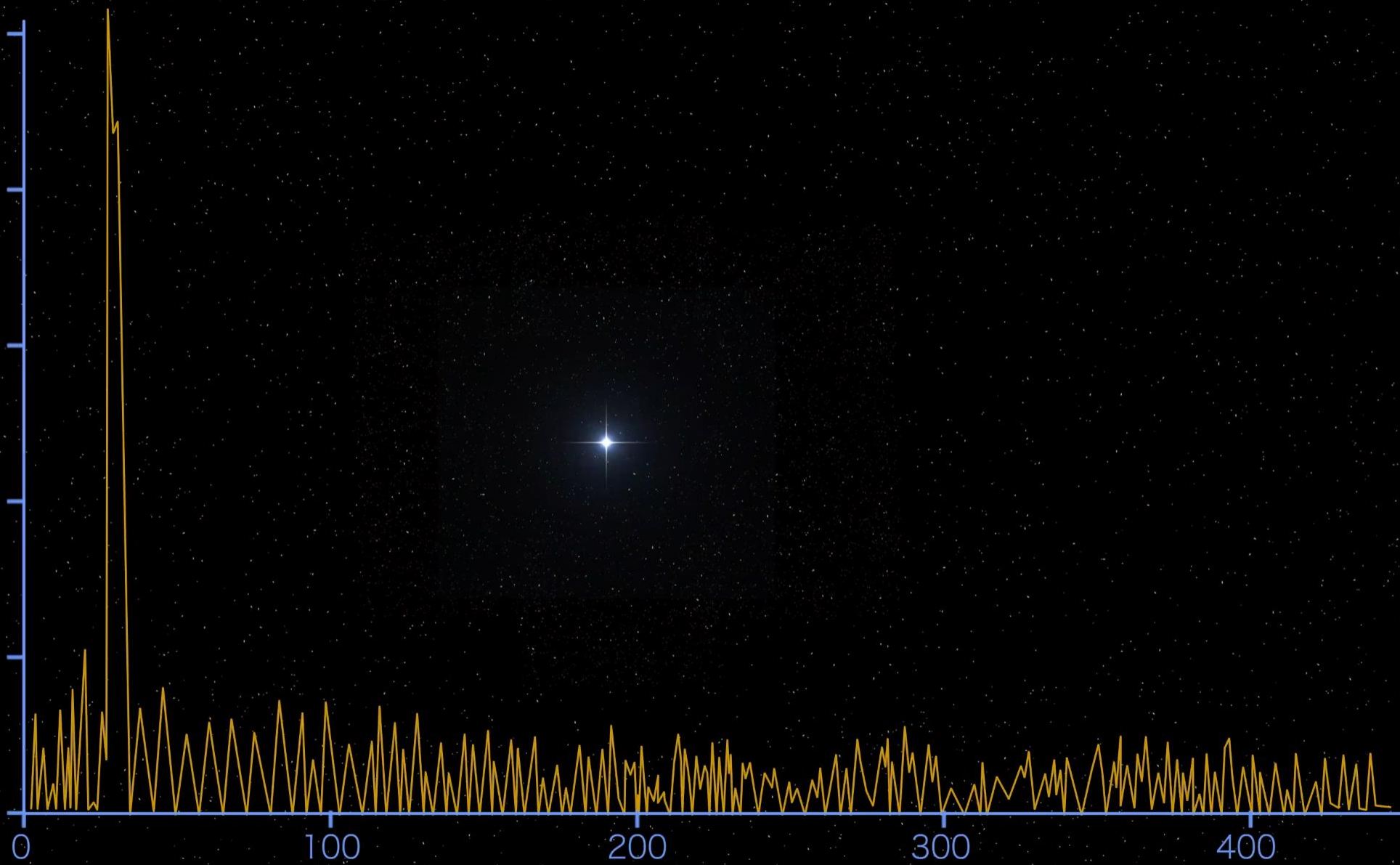
Roberts+ Nature 589 (2021)



## Magnetar Giant Flare



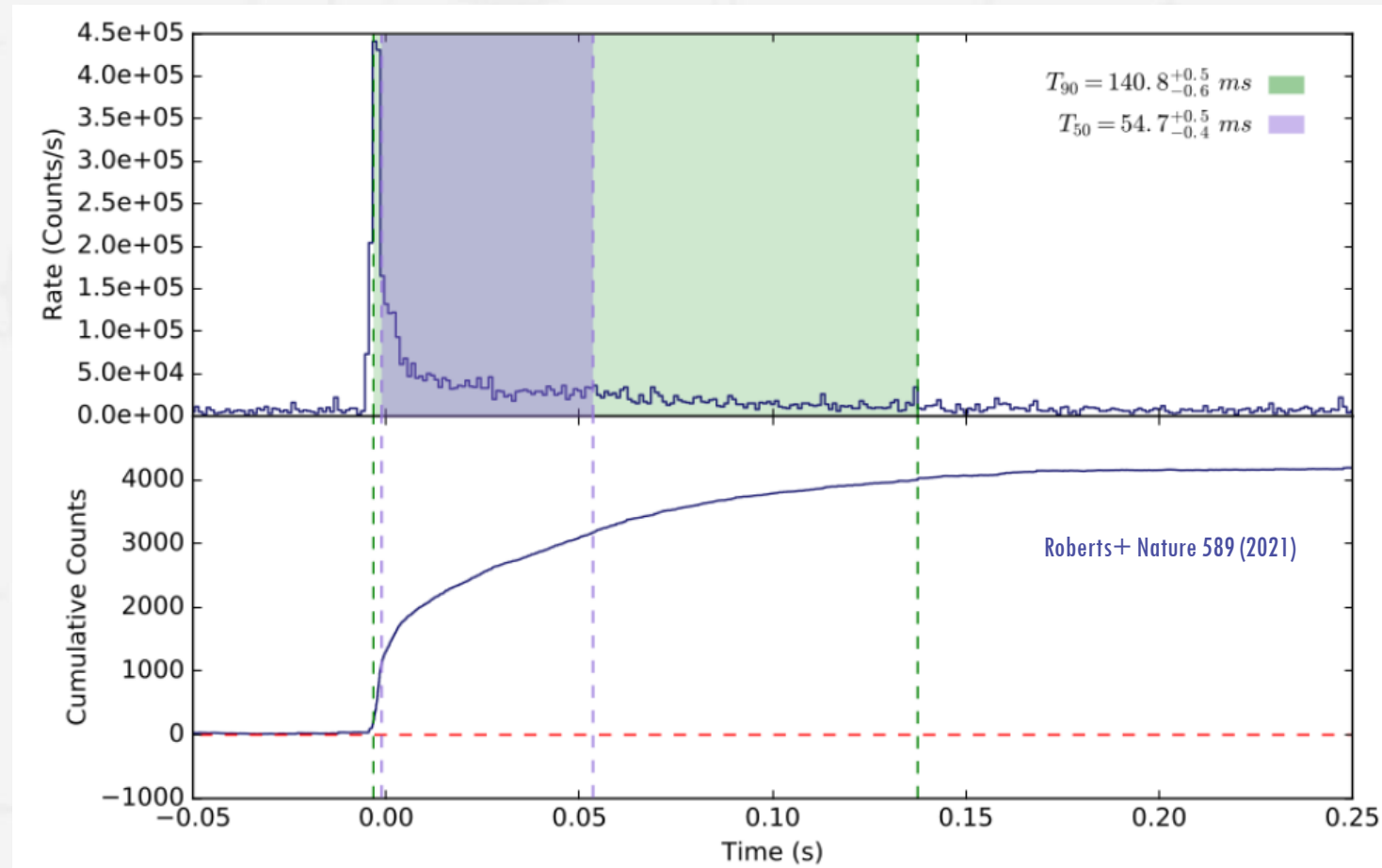
Photons  
per  
second



Time in seconds

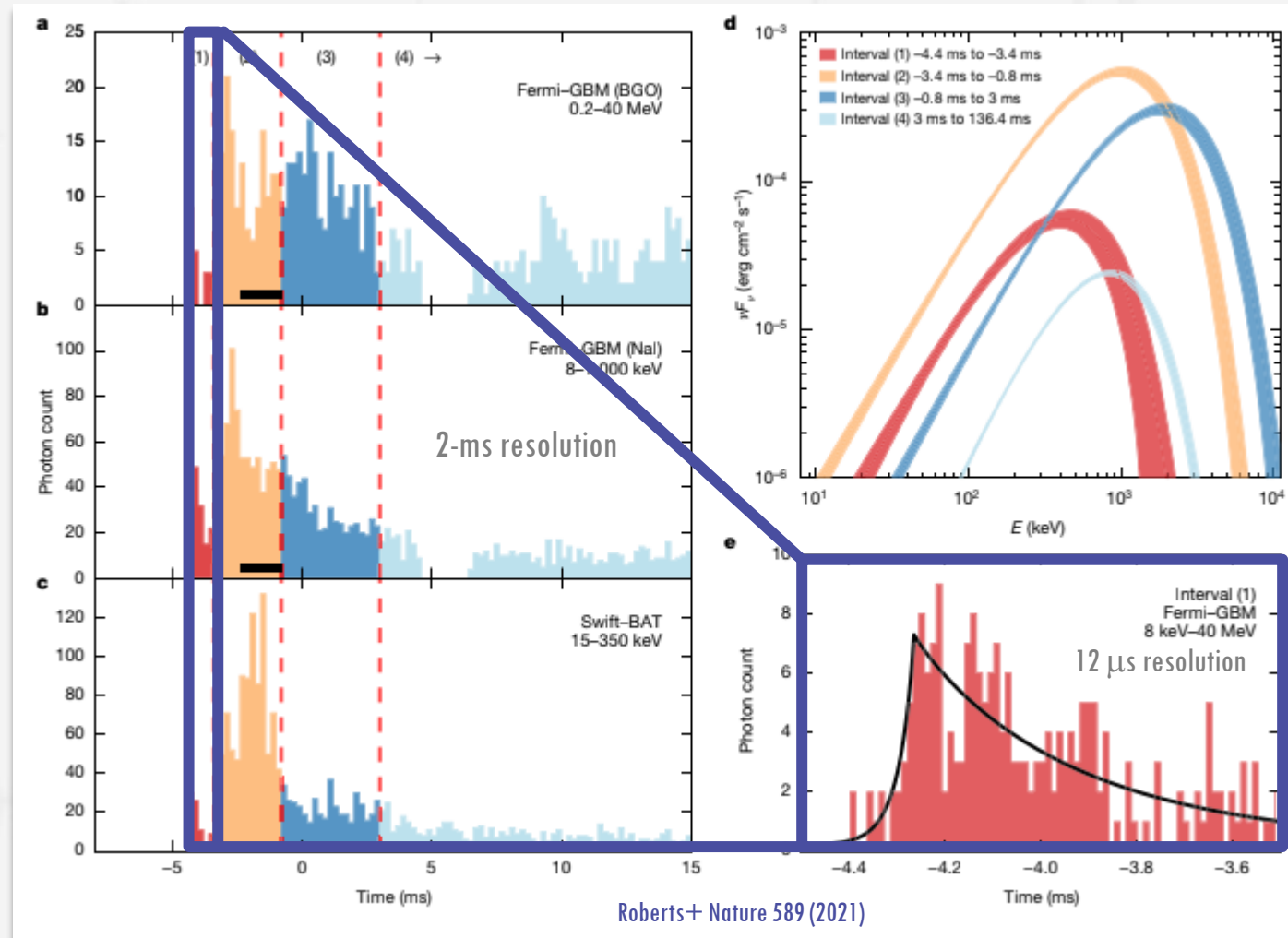
# GRB 200415A Fermi-GBM Observations

- The  $T_{90}$  (green) and  $T_{50}$  (purple) durations were calculated using the Swift–BAT data (15 – 350 keV) in count space. The errors are at the  $1\sigma$  confidence level.



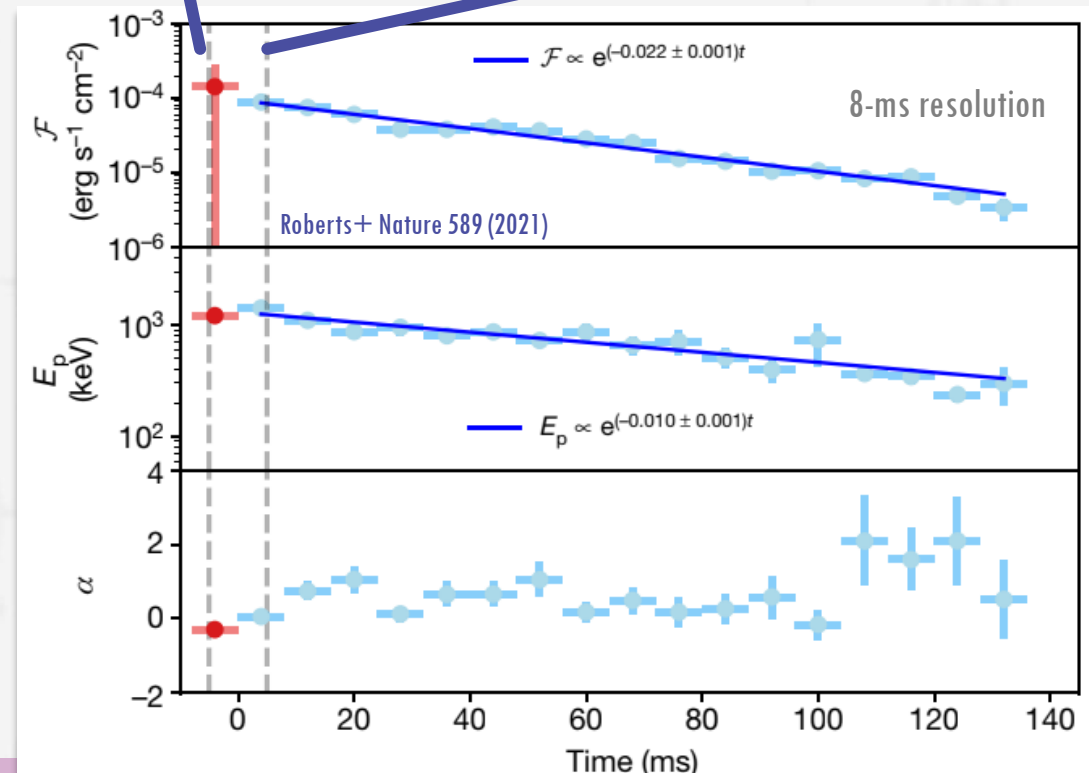
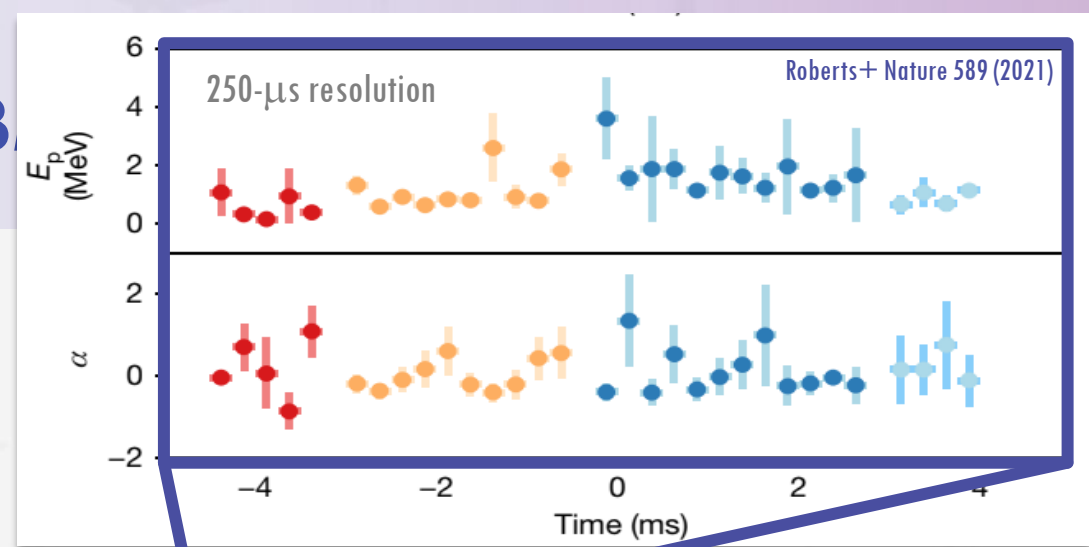
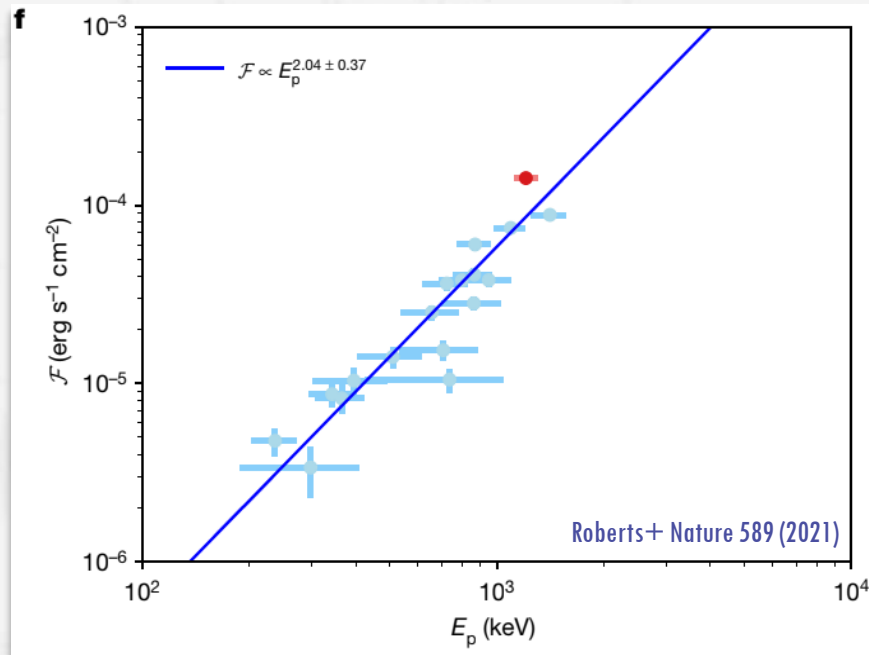


- Total energy:  $E_{\text{iso}} = 1.5 \times 10^{46}$  erg
- Total luminosity:  $L_{\text{iso}} = 1.1 \times 10^{47}$  erg s<sup>-1</sup>
- **Highest energy** GBM photon: **3 MeV**
- **Sub-ms variability** in photon energy
- Very rapid **flare risetime**  $\sim 77$   $\mu\text{s}$
- No instrumental limitations due to the source
  - **No modulated tail** after the initial 140 ms long bright burst
  - Possible signature of seismic vibrations  $\rightarrow$  Candidate **QPO** ( $2.5\sigma$ ) at 180 Hz
- **No radio** counterparts (VLA)

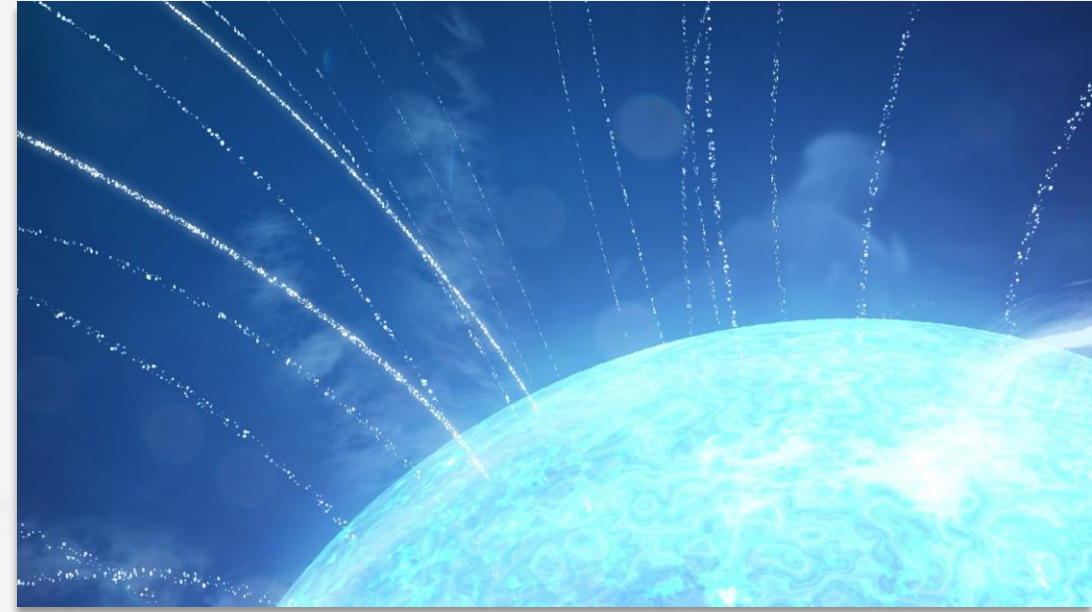


# GRB 200415A Fermi-GB

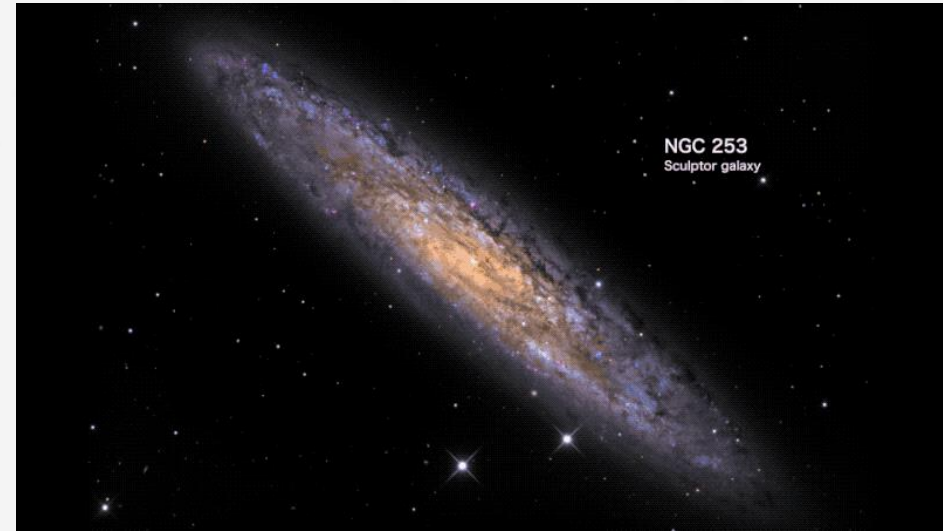
- Comptonized spectrum throughout, non-thermal
  - Flat spectral index ( $\alpha = 0$ ) over burst duration
  - Energy Flux Decay;  $\tau_{\mathcal{F}} = 45 \pm 3$  ms
  - Peak Energy Decay;  $\tau_{E_p} = 100 \pm 1$  ms
- Distinctive  $\mathcal{F} \propto E_p^2$  correlation



- The flux, and spectral shape of GRB 200415A are **unusual/unlikely for a short GRB**, when compared to catalogs from previous space missions
- Origin of MGFs: Energy release by **crustal fractures** ejects **hot plasma** into the inner magnetosphere
  - MeV-band emission must come from a **relativistic outflow** ( $>0.98c$ ) that is initially highly opaque
  - Spectral index  $\sim 0$ : Wind that is **highly opaque** to electron scattering (so-called “Compton cloud”)
  - **Inconsistent** with synchrotron GRB emission scenarios
- Sub-ms spectral evolution: **Relativistic lighthouse beaming effect** or relativistic-boosted wind acceleration and subsequent coasting/cooling
- 77  $\mu$ s risetime: **Extremely unlikely for a GRB**



- Can we confirm Flux and  $E_p$  decay trends,  $\text{Flux} \propto E_p^2$  ?
- Do we expect photon energies  $>3$  MeV?
- Are sub-ms spectral evolution and sub-ms risetime (light curve variability) expected in short cosmological GRBs too or is this “purely” a MGF phenomenon?
- What about the galaxies (SFR, metallicities, etc) that produce them?
  - Star-forming galaxies only?
- QPOs in the tails of galactic giant flares?
- Rate of giant flares per magnetar:  $< 0.02$  /yr – Repeating MGFs/sGRBs?
- Are 2% of all detected short GRBs in fact due to MGFs (0.3% of total GRBs)?
- Do all giant flares produce GeV emission?
- Is the plasma ultra-relativistic? (1806-20 was  $0.7c$ )
- Many more...



Credit: Copyright Adam Block/Mount Lemmon SkyCenter/University of Arizona

- 200415A is the clearest example yet of a **giant flare from an extragalactic magnetar**.
  - We need more event like this to link MGFs as potential source of sGRBs
- Unsaturated instrument spectra allow **new discoveries** to be made, opening a new front on studying the emission mechanism of these cataclysmic events
- These “repeating GRBs” may contribute a sizeable fraction of events to be present in GRB catalogs



Credit: Copyright Adam Block/Mount Lemmon SkyCenter/University of Arizona

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### Paper Links

**IPN:** <https://www.nature.com/ar@cles/s41586-020-03076-9>

**GBM/GUANO:** <http://doi.org/10.1038/s41586-020-03077-8>

**LAT:** <http://doi.org/10.1038/s41550-020-01287-8>

**Populations paper:** <https://iopscience.iop.org/ar@cle/10.3847/2041-8213/abd8c8>

### NASA press release

<https://www.nasa.gov/feature/goddard/2021/nasa-missions-unmask-magnetar-erup@ons-in-nearby-galaxies>

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**Thank you!**

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