

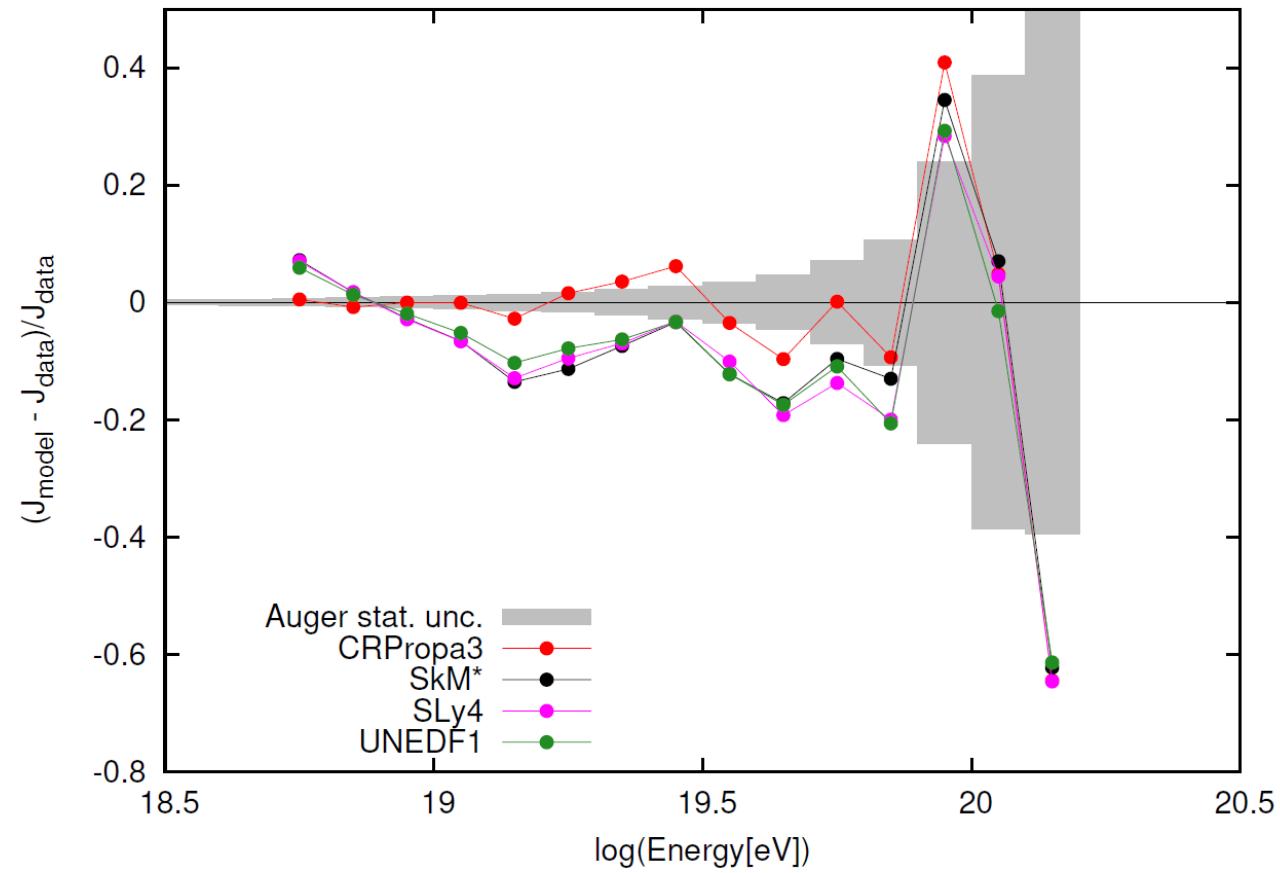
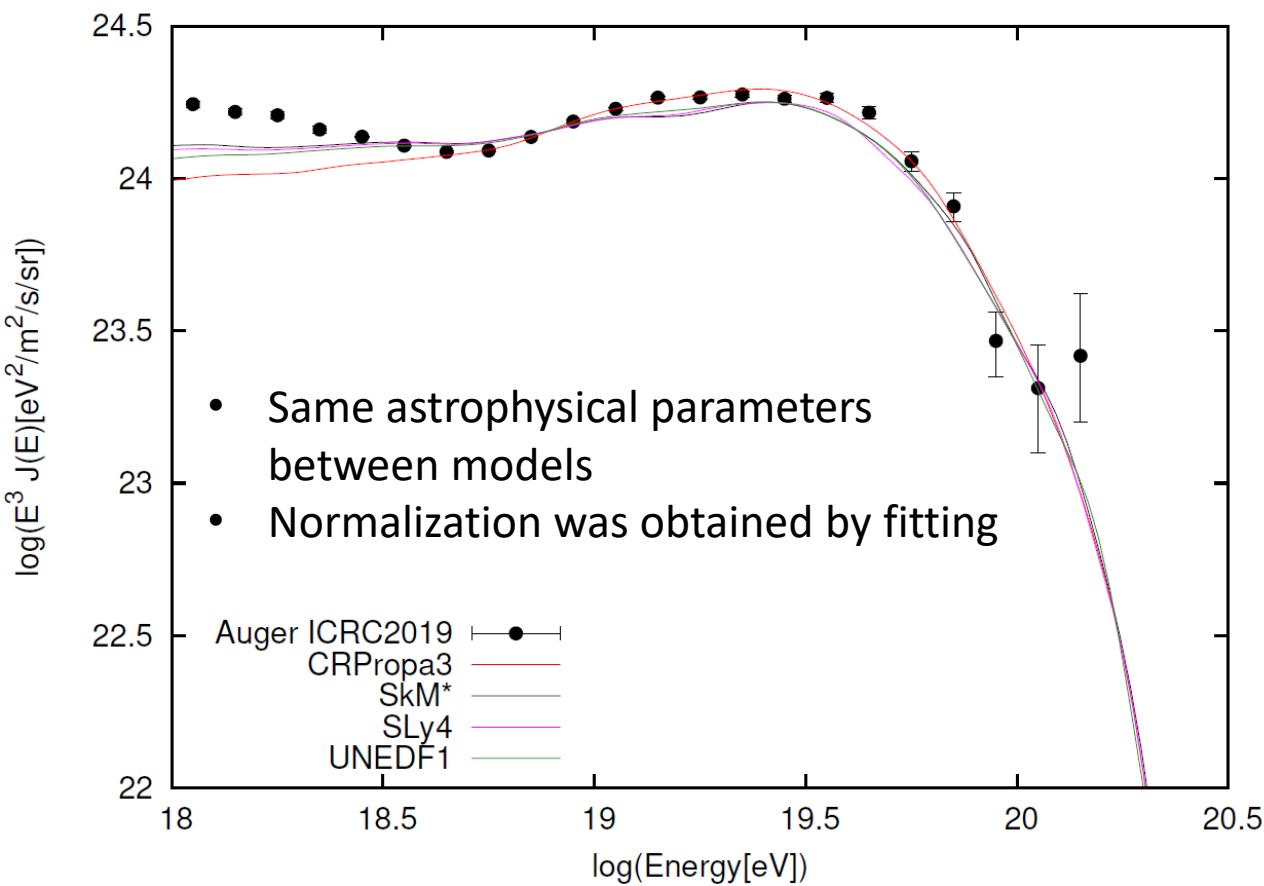
The impact of photonuclear reaction models on propagation of ultrahigh energy cosmic rays

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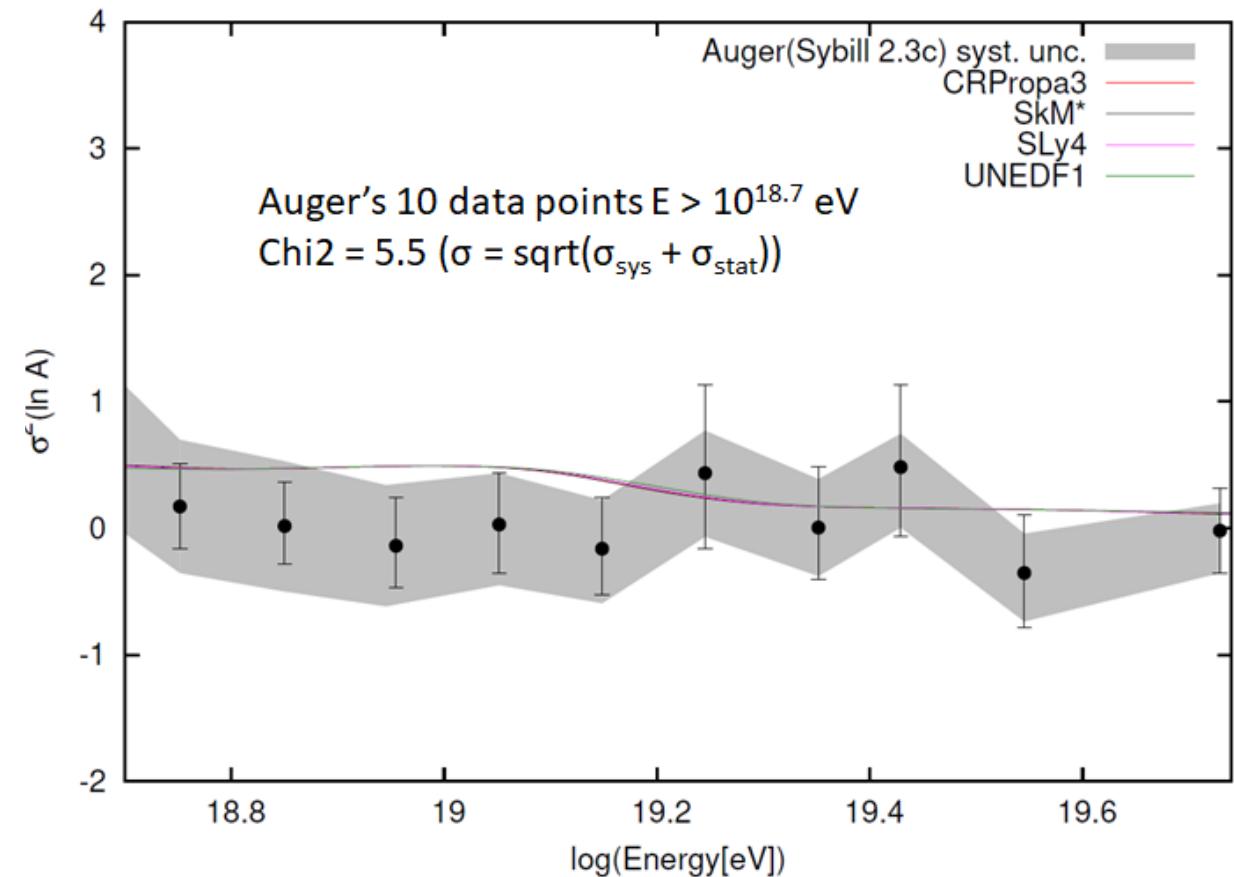
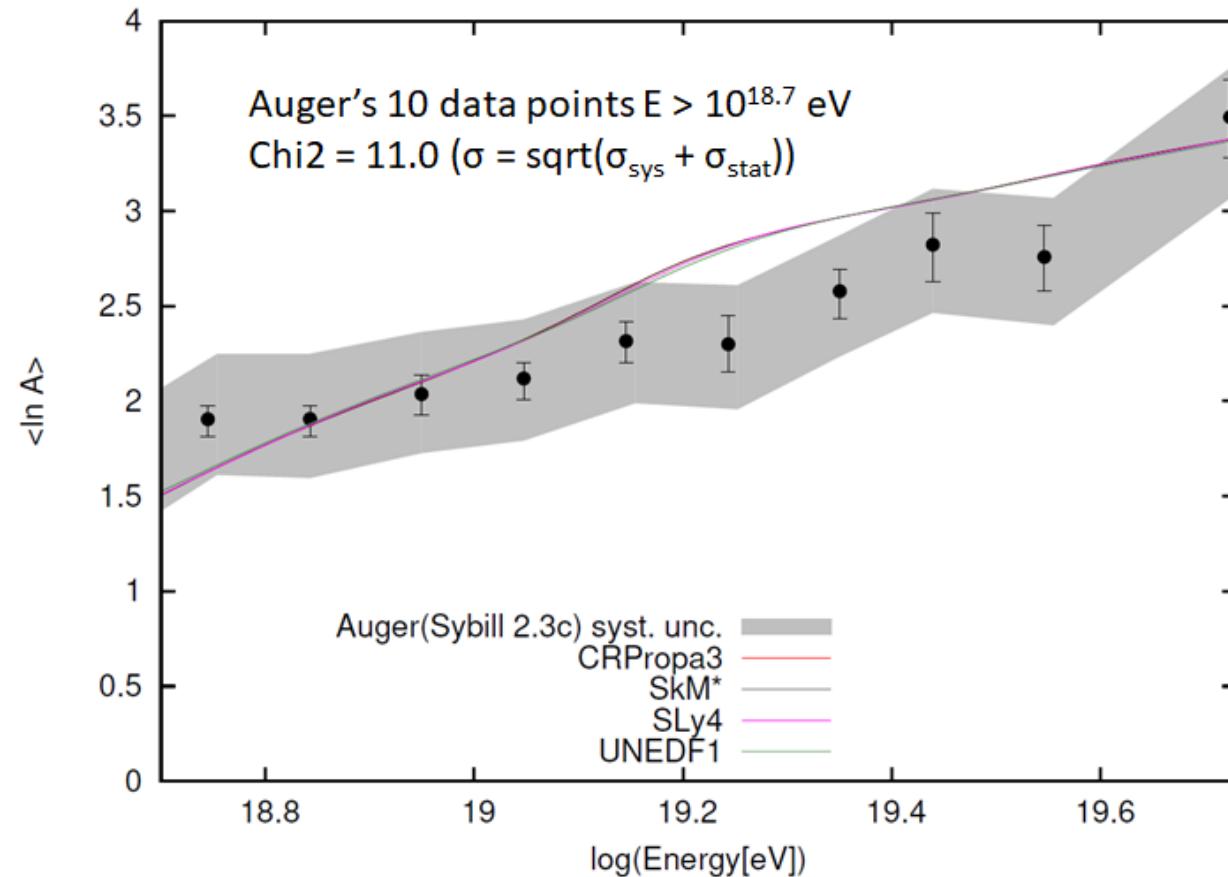
Comparison of Photonuclear Reaction Models in Simulations of Propagation of Ultrahigh Energy Cosmic Ray Nuclei using CRPropa

- T. Inakura *et al.*, *Phys. Rev. C* **80**, 044301 (2009)
and T. Inakura *et al.*, *Phys. Rev. C* **84**, 021302(R) (2011).
 - **The random phase approximation (RPA) calculations in Density Functional Theory (DFT)**
 - 3 interaction models
 - **SkM*** : J. Bartel *et al.*, *Nucl. Phys. A* **386**, 79 (1982).
 - **SLy4** : E. Chanbanat, P. Bonche, P. Haensel, J. Mayer, and R. Schaeffer, *Nucl. Phys. A* **627**, 710 (1997).
 - **UNEDF1** : M. Kortelainen *et al.*, *Phys. Rev. C* **85**, 024304 (2012).
- TALYS https://tendl.web.psi.ch/tendl_2019/talys.html
 - Included in **CRPropa**
 - Statistical Hauser-Feshbach theory etc.

Spectrum Fit (Auger ICRC2019)



$\langle \ln A \rangle$, $\sigma^2(\ln A)$ (Auger ICRC2019)



Summary

- We implemented the random phase approximation (**RPA**) calculations in density functional theory (**DFT**) (T. Inakura *et al.*) in CRPropa and simulated propagation of ultrahigh energy cosmic ray nuclei.
- We found that the difference between the RPA calculations and the default settings of CRPropa in the spectral shape is much larger than the statistical uncertainty of the experimental data when the same astrophysical parameters are assumed. The model predictions of photonuclear reactions will be experimentally tested by the **PANDORA** project in the near future.