



# Forbush decrease on September 6-13, 2017 observed by the Tanca water-Cherenkov detector

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# Introduction

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- September of 2017: numerous solar events were registered, among them, three flares with intensity M5.5, X9.3 and X8.2 that originated three halo CMEs (Coronal Mass Ejection). [1,2]

**Table 1:** Halo CMEs registered by LASCO/C2 in September 2017.

	<b>Arrival Time</b>	<b>Max Kp</b>	<b>Dst min.(nT)</b>	<b>Dst min. Time</b>	<b>Origin</b>
<i>CME1</i>	2017-09-06T23:08Z	4	-23	2017-09-07T09:00Z	M5.5 flare
<i>CME2</i>	2017-09-07T22:30Z	8	-142	2017-09-08T02:00Z	X9.3 flare
<i>CME3</i>	2017-09-12T19:26Z	5	-50	2017-09-13T01:00Z	X8.2 flare

# Tanca

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- The CMEs produced Forbush decreases on the signal of a muon detector located at the University of Campinas, Brazil, called Tanca (*Tanque de Campinas*).
- Tanca is a replica of the detectors used in Pierre Auger Observatory and it is part of the LAGO (Latin Giant Observatory) collaboration.



Figure 1. Tanca.

# Results

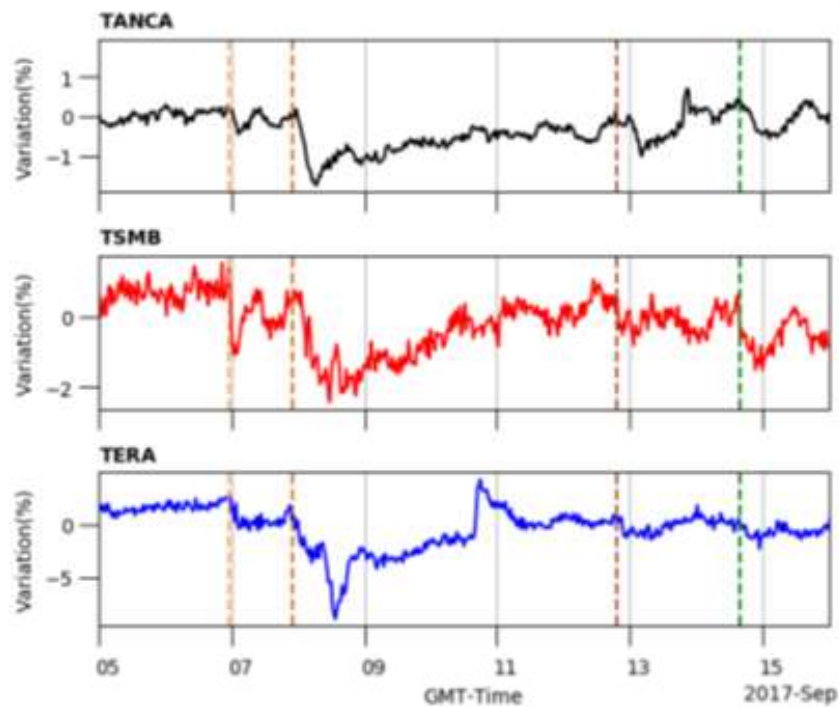


Figure 2. Cosmic ray signal from Tanca ( $R_c = 9.36$  GV) in black, Tsmb ( $R_c = 9.15$  GV) in red and Tera ( $R_c = 0.01$  GV) in September 2017. Tsmb and Tera data are available in Reference [3].

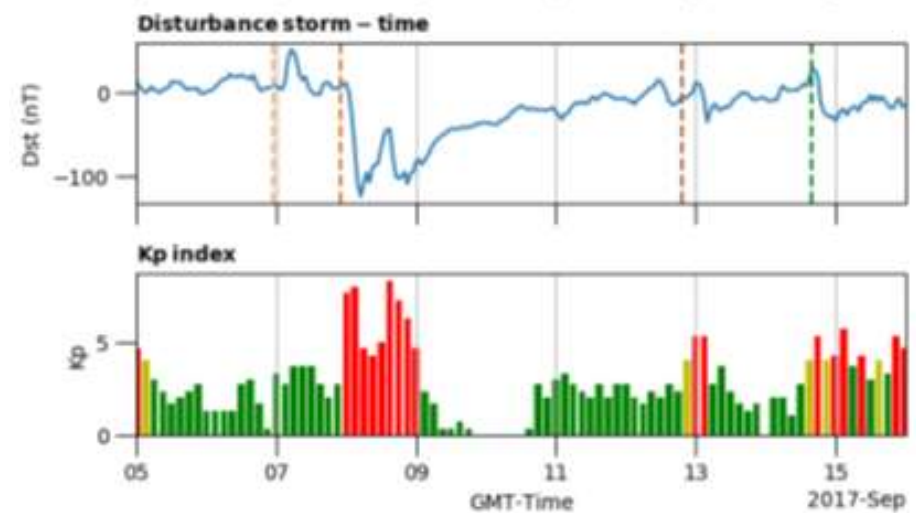


Figure 3. Hourly averaged Dst in cyan and the Kp index (green:  $K_p < 4$ , yellow:  $K_p = 4$  and red:  $K_p > 4$ ) in September 2017. [4]

# Conclusion

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- Tanca registered four signal decreases: three caused by the shock of CMEs in arrival time at 2017-09-06, 2017-09-07 and 2017-09-12 and one by a stream interaction region in 2017-09-14
- These observations of Forbush events carried out by the Tanca detector show the effects of solar events on the Earth's magnetic field for a region of energy above the energies observed by space missions, adding information about the effects of the more energetic particles of these events.

# Acknowledgements

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# References

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- [1] SOHO LASCO CME CATALOG - CDAW DATA CENTER. url: [https://cdaw.gsfc.nasa.gov/CME\\_list/](https://cdaw.gsfc.nasa.gov/CME_list/)
- [2] CME Scoreboard. url: <https://kauai.ccmc.gsfc.nasa.gov/CMEscoreboard/>.
- [3] Real-Time Database for high-resolution Neutron Monitor measurements. url: <http://www01.nmdb.eu/>.
- [4] World Data Center for Geomagnetism, Kyoto. url: <http://wdc.kugi.kyoto-u.ac.jp/index.html>.
- [5] Jingnan Guo et al. “Modeling the Evolution and Propagation of 10 September 2017 CMEs and SEPs Arriving at Mars Constrained by Remote Sensing and In Situ Measurement”. In: *Space Weather* 16.8 (2018), pp. 1156–1169. doi: 10.1029/2018sw001973
- [6] S. E. Forbush. “On the Effects in Cosmic-Ray Intensity Observed During the Recent Magnetic Storm”. In: *Physical Review* 51.12 (1937), pp. 1108–1109. doi: 10.1103/physrev.51.1108.3.
- [7] I. M. Chertok, A. V. Belov, and A. A. Abunin. “Solar Eruptions, Forbush Decreases, and Geomagnetic Disturbances From Outstanding Active Region 12673”. In: *Space Weather* 16.10 (2018), pp. 1549–1560. doi: 10.1029/2018sw001899.
- [8] I. Sidelnik [LAGO], PoS ICRC2015(2016), 665 doi:10.22323/1.236.0665
- [9] Cutoff Rigidity Calculator. url: <http://cosmos.hwr.arizona.edu/Util/computeCR.php>.
- [10] CERCLÉ. url: <https://previ.obspm.fr/>

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Thank you!