

ISS-CREAM detector performance and tracking algorithms

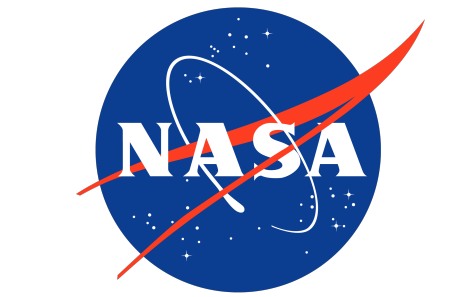


Kenichi Sakai^{a,b}, Scott L. Nutter^c, Tyler Anderson^d, Yu Chen^d, Stephane Coutu^d, Tyler LaBree^c, Jason T. Link^{a,b}, John W. Mitchell^b, S. A. Isaac Mognet^d, Jacob Smith^{a,b}, Monong Yu^d



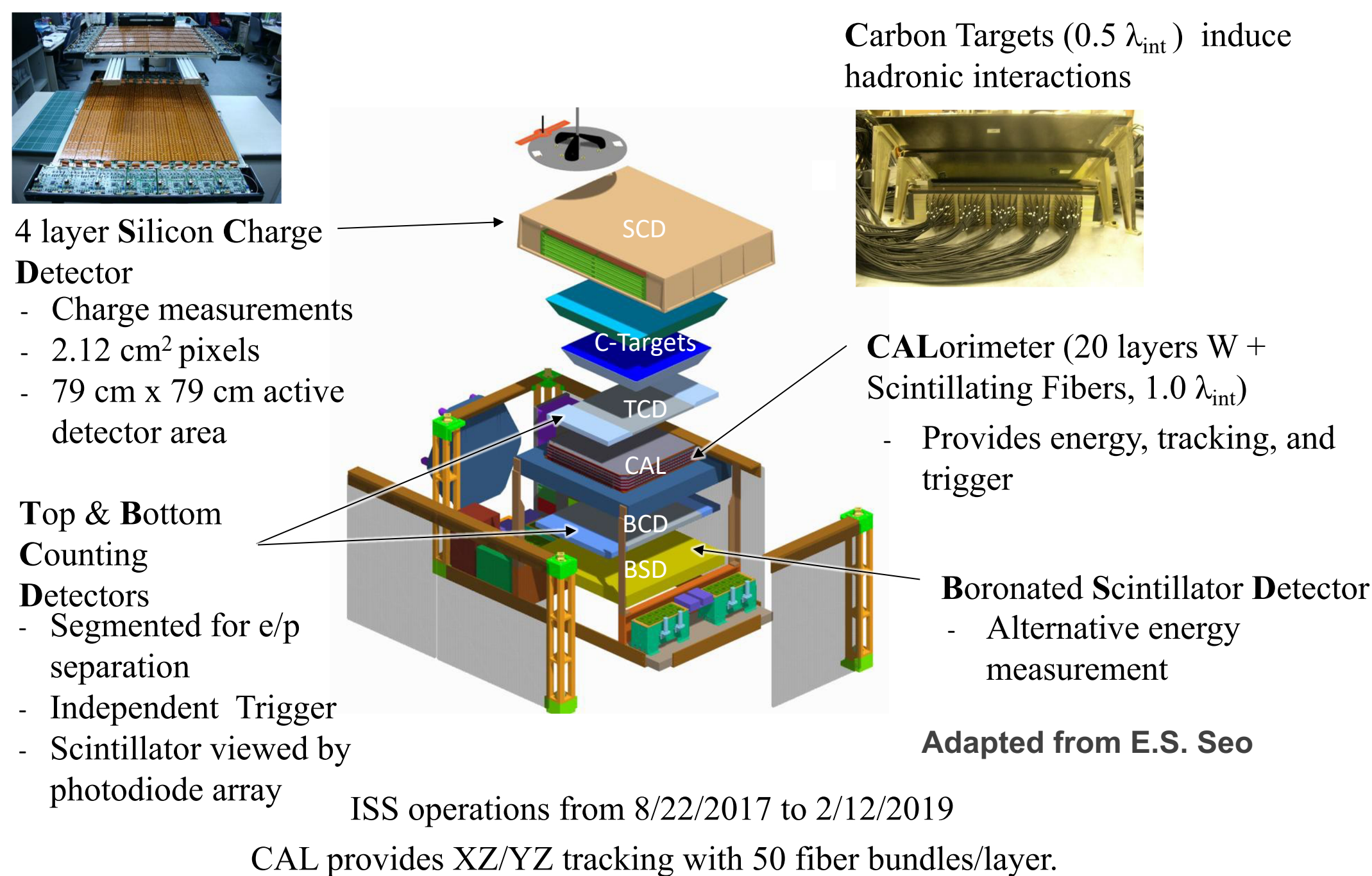
* Presenter: kenichi.sakai@nasa.gov

a. Center for Research and Exploration in Space Science and Technology (CRESST), UMBC, Baltimore MD, 21250 USA
b. NASA Goddard Space Flight Center, Astroparticle Physics Laboratory, Greenbelt, MD 20771, USA
c. Northern Kentucky University, Dept of Physics, Geology, and Engineering Technology, Highland Heights, KY 41099 USA
d. Penn State University, Department of Physics, University Park, PA 16802, USA



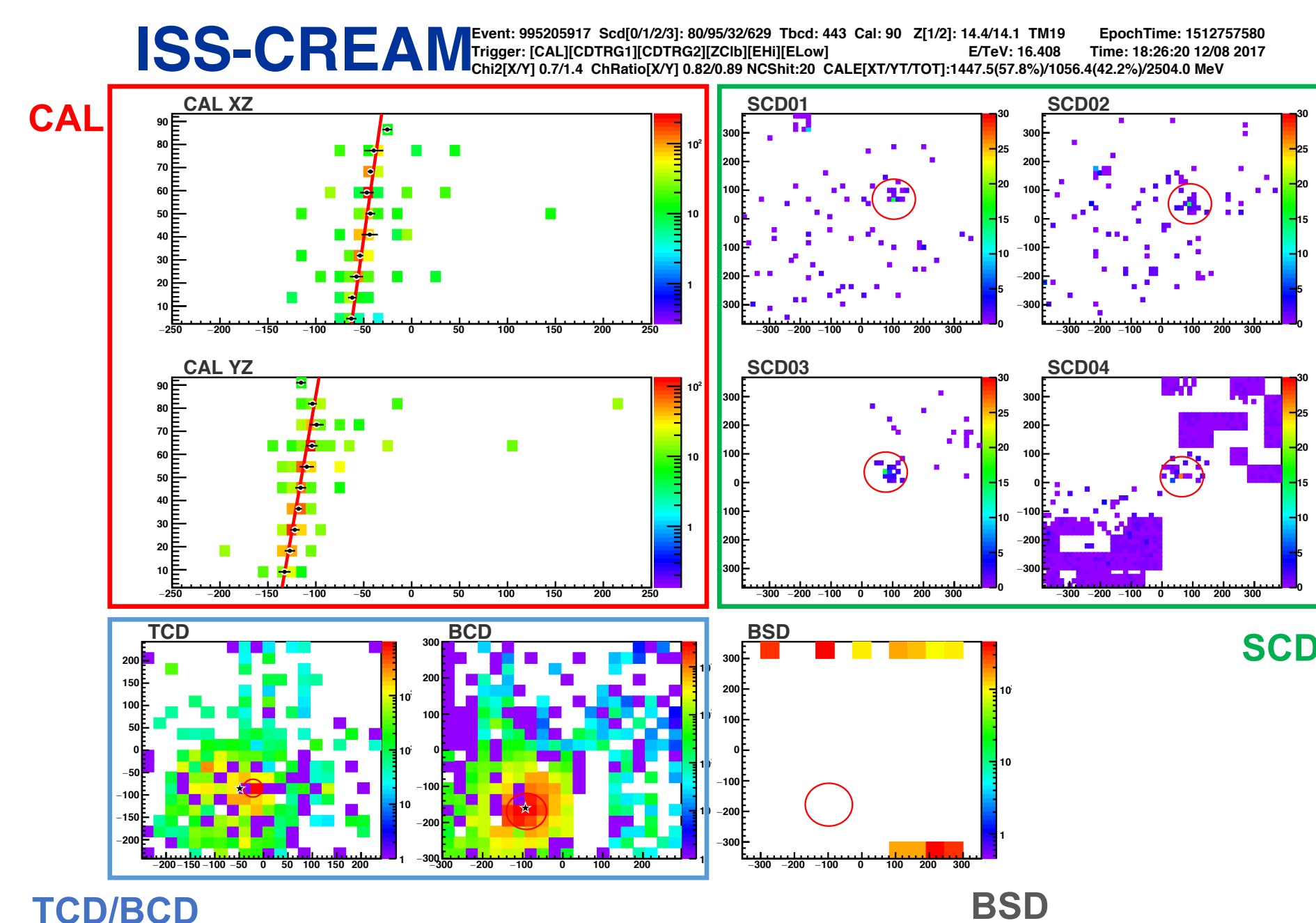
The ISS-CREAM experiment grew out of a balloon-borne instrument, CREAM, which measures the fluxes of elements between protons and iron in an energy region almost up to the knee.

1. ISS-CREAM instrument

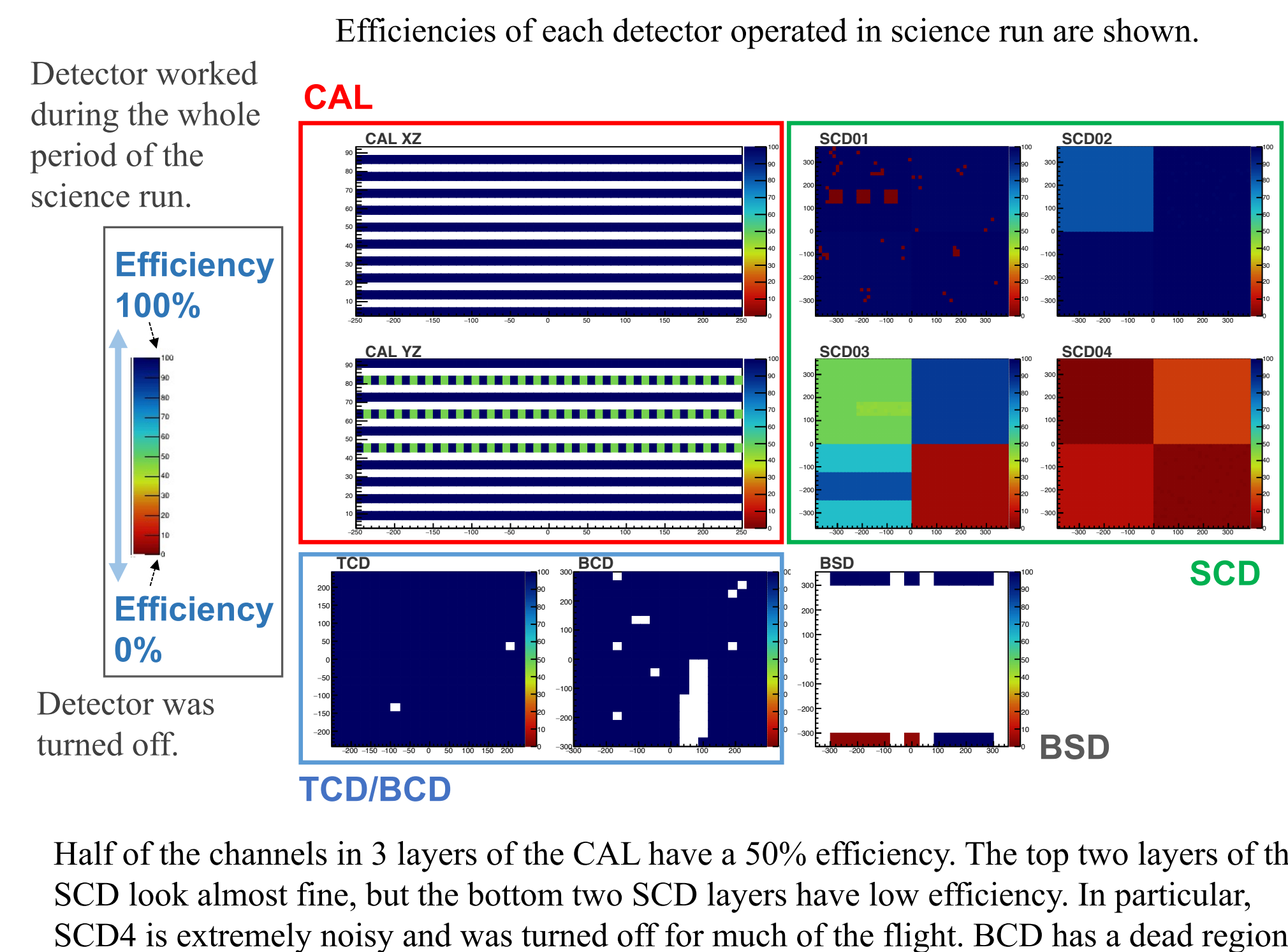


2. Event display

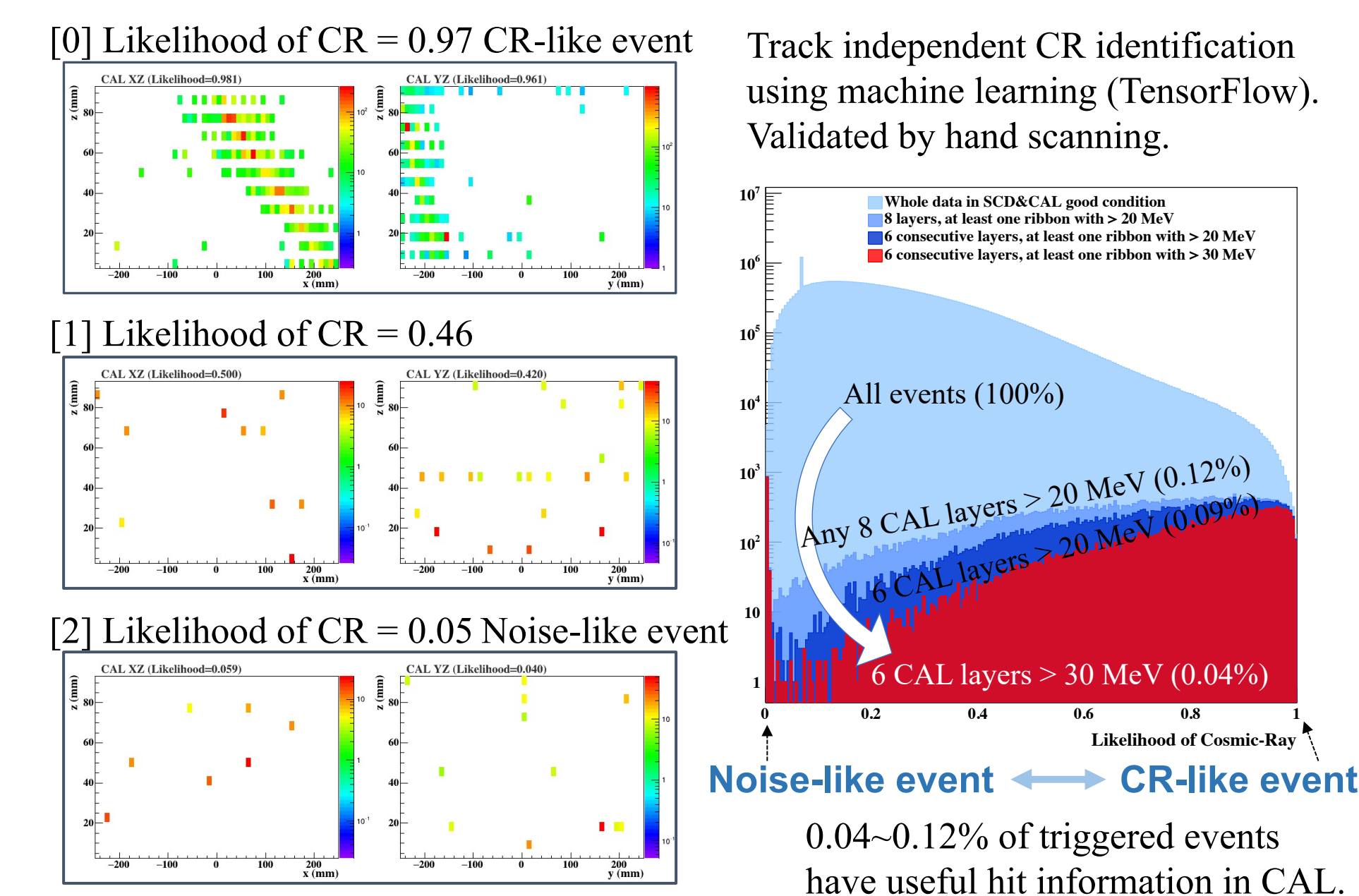
Event display enables simultaneous viewing of all detector responses.
Shown: Typical high energy event. Charge: Si (Z=14)



3. Detection efficiencies of each detector during the observations



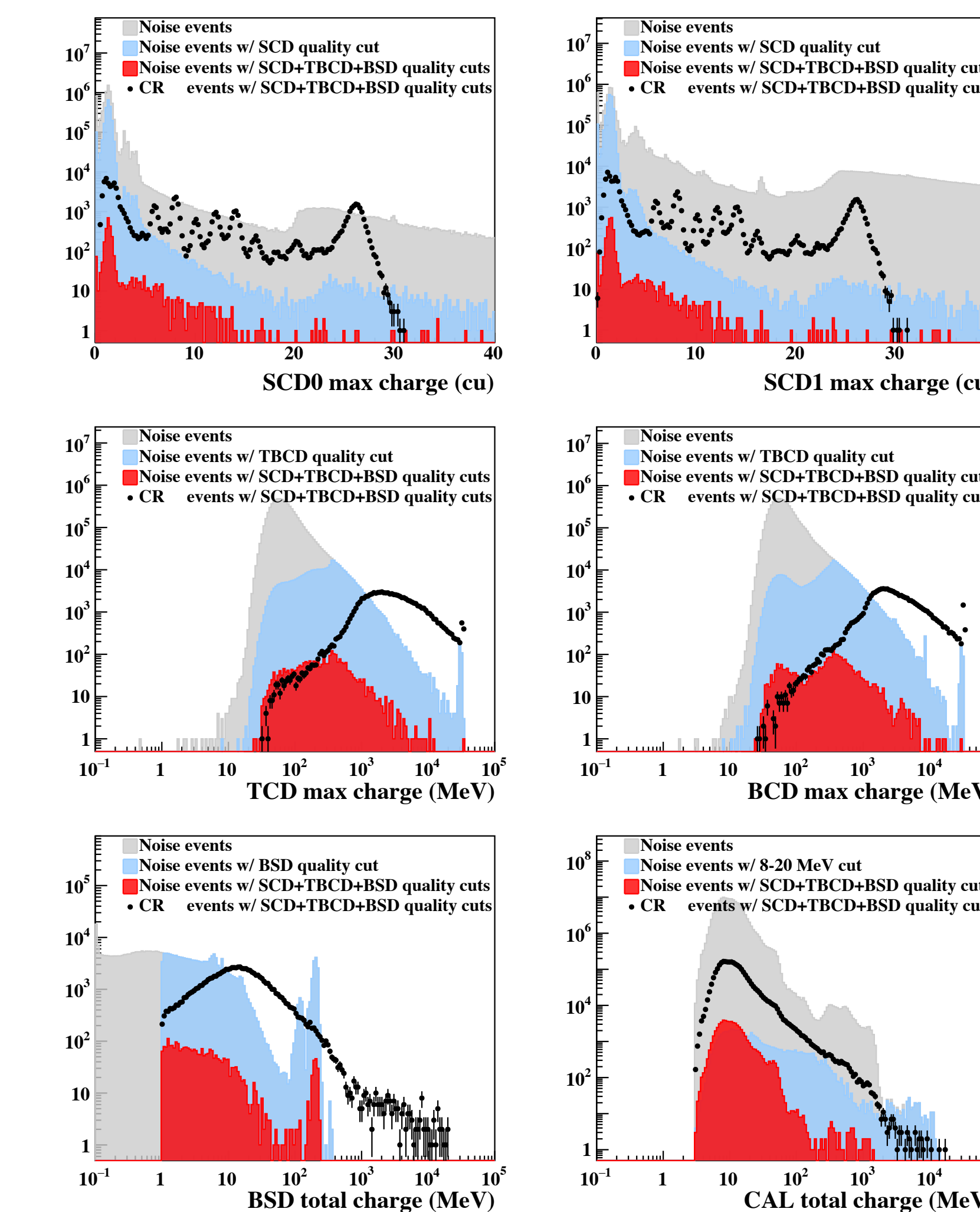
5. Detector performance of the CAL



• Detail of machine learning → ICRC Poster 476 by Monong Yu
• Energy calibration by using BSD → ICRC Poster 866 by Yu Chen

4. Detector noise characterization

The background detector noise characterization was investigated by using 0.5 Hz periodic trigger that has identical acquisition process to event data.



Above figures show that the combined quality cuts of SCD, T/BCD and BSD worked well to prevent the contamination by noise events. The reason that CAL quality cut isn't applied is described in Section 5.

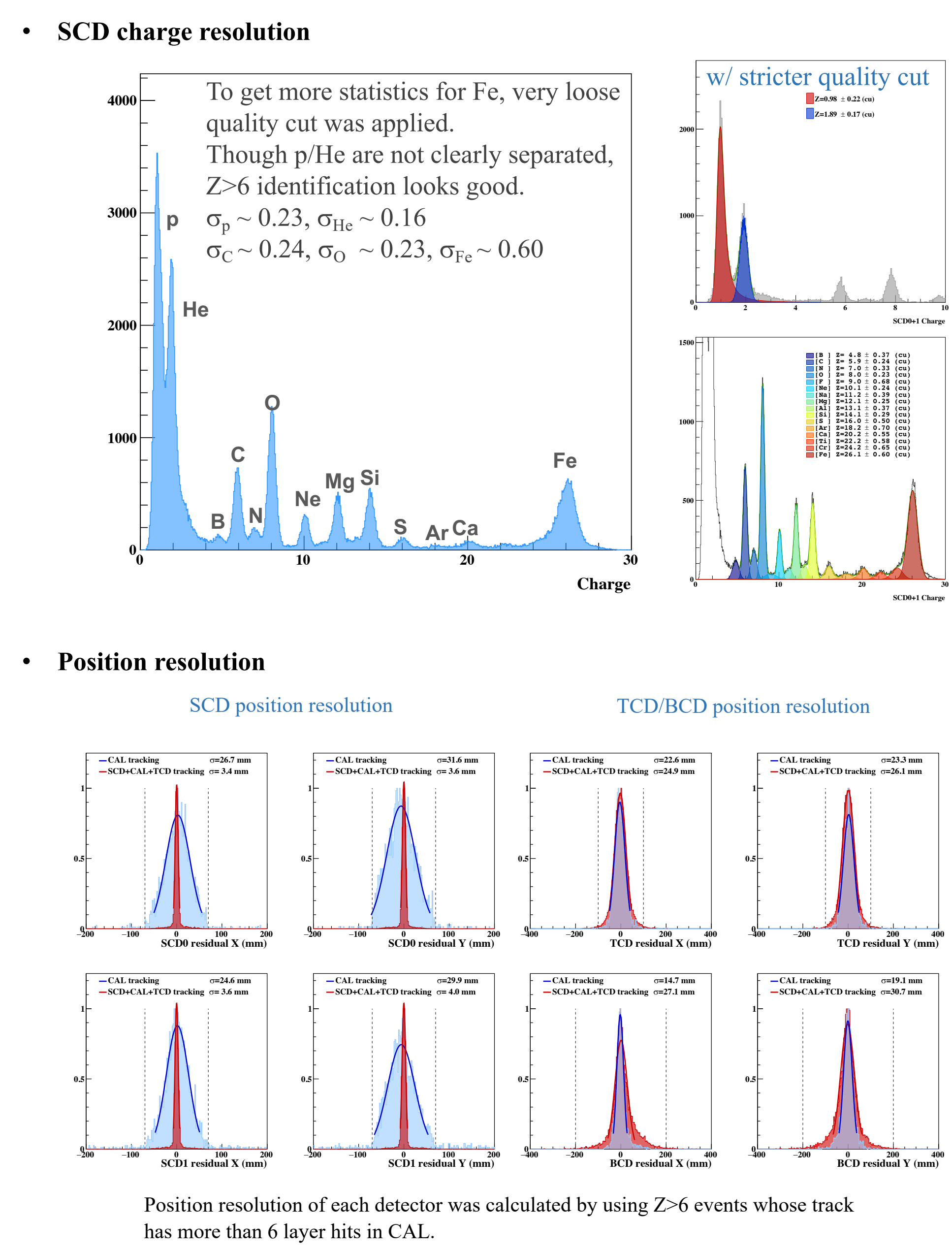
6. Tracking algorithms (22)

The following two are picked up from 22 different tracking algorithms.

CAL-based tracking: standard tracking: track was reconstructed by using energy-weighted CAL hits, then SCD charge is searched within a 70mm radius of the extrapolated point. For most events, CAL tracking failed because more than 99.9% events don't have useful hit information in the CAL.

SCD/CAL/TCD-based tracking: new tracking using 2 layers of SCD, TCD and CAL hits was developed and used for the flux calculation. Even if CAL is missing data, the tracking was reconstructed by using three points from the other detectors.

7. Detector performance of the SCD and T/BCD



• Flux calculation → ICRC Poster 696 by Scott L. Nutter

Acknowledgements

This work was supported in the U.S. by NASA grants NNX17AB43G, NNX17AB42G, and their predecessor grants, as well as by directed RTOF funds to NASA GSFC. The authors also thank M. Geske, Penn State, for contributions to the BSD, and K. Wallace at Northern Kentucky University for contributions to Monte Carlo simulations. We also recognize the contributions of past CREAM and ISS-CREAM collaborators.