

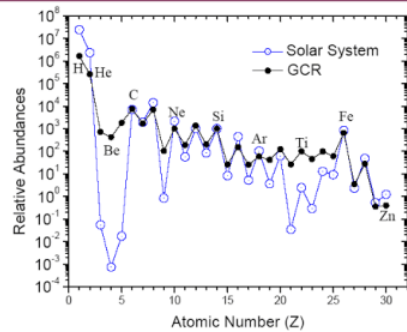
# A Data-Driven approach for the measurement of $^{10}\text{Be}/^9\text{Be}$ flux ratio in Cosmic Rays with magnetic spectrometers

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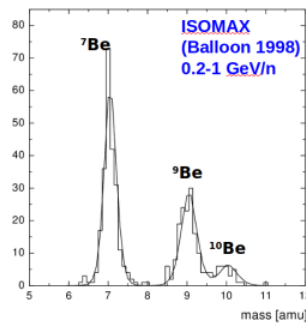
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The  $^{10}\text{Be}/^9\text{Be}$  flux ratio (thanks to the 2 My lifetime of  $^{10}\text{Be}$ ) is a radioactive clock providing the measurement of CR residence time in the Galaxy. Existing measurements of  $^{10}\text{Be}/^9\text{Be}$  in CR are limited to low energy and affected by large uncertainties, in particular from the Montecarlo simulation. A Data-Driven approach in magnetic spectrometers is presented, as an example it is applied to PAMELA data providing a new measurement in the 0.25-0.85 GeV/n range.

## Isotopes of Beryllium in cosmic rays



Be is scarce in CR not produced by stellar nucleosynthesis  
Be is a secondary CR produced by CNO spallation with ISM



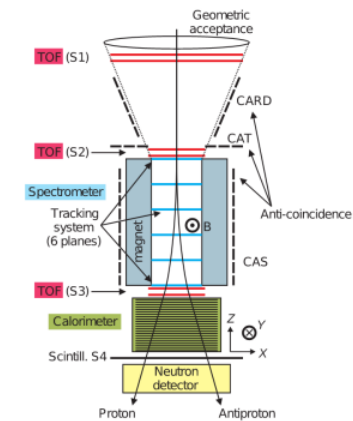
Be isotopes in CR:

$^7\text{Be}$  (E.C.  $T_{1/2}=55\text{d}$ )  
("naked"  $^7\text{Be}$  is stable)

$^9\text{Be}$  (stable)

$^{10}\text{Be}$  ( $\beta$   $T_{1/2}=1.39\text{My}$ )  
(is a radioactive clock)

## PAMELA experiment

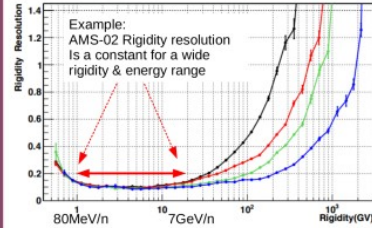


ToF:  
 $\beta = \frac{v}{c}$

TRK:  
 $\frac{RZ}{\beta\gamma} = M$

Calo:  
 $\frac{dE}{dX}(\beta)$

## Mass & R resolution



$$\frac{\Delta M}{M} = \sqrt{\left(\frac{\delta R}{R}\right)^2 + \gamma^4 \left(\frac{\delta \beta}{\beta}\right)^2}$$

For a fixed  $E_k/n = (\gamma - 1)0.9315 \text{ GeV}$

$\frac{\Delta M}{M} = \text{const.} \Rightarrow$  mass template scaling  
 Templates  $T_7$ ,  $T_9$  and  $T_{10}$  are the three unknown mass distributions.

## The “Data-Driven” approach: recipe summary

The three Be mass are similar, a linear approximation is applied.  $\sigma_a$  is the RMS of  $T_a$  and  $x_a$  is the median of  $T_a$ . The (linear) transformation  $L_{a,b}T_a = T_b$  is the function:  $x \rightarrow \frac{\sigma_a}{\sigma_b}x + [x_a - \frac{\sigma_a}{\sigma_b}x_b]$ .

The same transformation but applied to a different, template  $L_{a,b}T_c = T_d$  is:

$$\sigma_d = \sigma_c \frac{\sigma_b}{\sigma_a} \text{ and } x_d = x_b + (x_c - x_a) \frac{\sigma_b}{\sigma_a}$$

The known (measured) data distribution is  $D(x)$ , thus this system must be solved:

$$D(x) = {}^7\text{Be}T_7 + {}^9\text{Be}T_9 + {}^{10}\text{Be}T_{10}$$

$$L_{7,9}D(x) = {}^7\text{Be}T_9 + {}^9\text{Be}L_{7,9}T_9 + {}^{10}\text{Be}L_{7,9}T_{10}$$

$$L_{7,10}D(x) = {}^7\text{Be}T_{10} + {}^9\text{Be}L_{7,10}T_9 + {}^{10}\text{Be}L_{7,10}T_{10}$$

The  ${}^7\text{Be}$  template can be written as:

$$T_7 = \frac{1}{\sqrt{{}^7\text{Be}}} \left[ D - \frac{{}^9\text{Be}}{{}^7\text{Be}} L_{7,9}D - \frac{{}^{10}\text{Be}}{{}^7\text{Be}} L_{7,10}D \right] + \left( \frac{{}^9\text{Be}}{{}^7\text{Be}} \right)^2 T_{G1} + \frac{{}^{10}\text{Be}}{{}^7\text{Be}} \frac{{}^{10}\text{Be}}{{}^7\text{Be}} (T_{G2} + T_{G3}) + \left( \frac{{}^{10}\text{Be}}{{}^7\text{Be}} \right)^2 T_{G4}$$

the last four terms, are defined by:

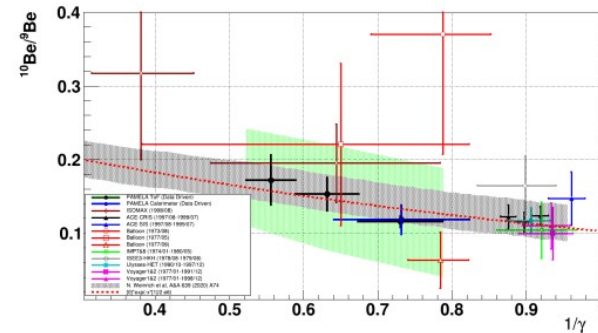
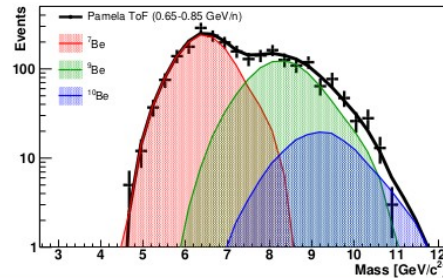
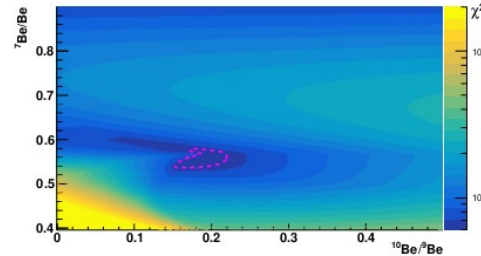
$$T_{G1} = L_{7,9}T_9 = L_{7,x_{G1}}T_7 \quad T_{G2} = L_{7,9}T_{10} = L_{7,x_{G2}}T_7$$

$$T_{G3} = L_{7,10}T_9 = L_{7,x_{G3}}T_7 \quad T_{G4} = L_{7,10}T_{10} = L_{7,x_{G4}}T_7$$

$T_7$  can be iteratively evaluated for each fixed  ${}^7\text{Be} > {}^9\text{Be} > {}^{10}\text{Be}$  configuration

$T_9$  and  $T_{10}$  are obtained by scaling  $T_7$  and a  $\chi^2$  is evaluated. Three un-physical  $\chi^2 = 0$  solutions are  ${}^n\text{Be}/{}^m\text{Be}=1$ . Use of the statistical bootstrap is suggested for confidence intervals of physical solution.

## “Data-Driven” applied to PAMELA Be events: a new ${}^{10}\text{Be}/{}^9\text{Be}$ measurement



Data-Driven approach allows a new measurement in 0.25-0.85 GeV/n.

Green shaded area is a (cautious) systematic error.

First experimental hint for time dilation effect in  ${}^{10}\text{Be}/{}^9\text{Be}$ .

Adopting a minimal model:  ${}^{10}\text{Be}/{}^9\text{Be} = Ae^{-\frac{T}{\tau}}$  (known  $\tau = 2\text{My}$ )

$A = 0.27 \pm 0.13$  and  $T = 1.9 \pm 1.1 \text{ My}$  (dominated by PAMELA data)

Data-Driven approach allows an independent test of Montecarlo systematics. Next step is the measurement using AMS-02 data.